

PHYS4150 — PLASMA PHYSICS

LECTURE 11 -

*Sascha Kempf**

G135, University of Colorado, Boulder

Fall 2020

1 FLUID DESCRIPTION OF PLASMAS

1.1 *Fluid variables*

We first define the volume of a fluid parcel as

$$dx^3 = dx dy dz,$$

which has the mass

$$m = (n_i m_i + n_e m_e) dx^3$$

and the mass density

$$\rho = \frac{m}{dx^3} = (n_i m_i + n_e m_e) = \rho_i + \rho_e.$$

Mass m and density ρ are fluid variables. We now introduce the average velocity \mathbf{u} of the plasma particles in the fluid parcel. Obviously, the fluid parcel will flow at this speed. The distribution $f(\mathbf{v}|\mathbf{u})$ of particle speeds can often be described by a shifted Maxwellian, i.e.

$$f(\mathbf{v}|\mathbf{u}) = \frac{n}{(\pi v_{th}^2)^{3/2}} \exp \left\{ -\frac{(\mathbf{v} - \mathbf{u})^2}{v_{th}^2} \right\}.$$

Knowledge of $f(\mathbf{v})$ allows us to compute \mathbf{u}

$$\mathbf{u} = \frac{\int \mathbf{v} f(\mathbf{v}) dx^3}{\int f(\mathbf{v}) dx^3} = \frac{1}{n} \int \mathbf{v} f(\mathbf{v}) dx^3.$$

*sascha.kempf@colorado.edu

The next fluid variable to discuss is the particle flux through a face of the fluid volume dx^3 . Let us consider the flux through the $dx dy$ face. Because flux is number of particles per area and time we can write

$$\Gamma_{xy} = \frac{N}{dx dy dt},$$

or after multiplying with dz/dz

$$\Gamma_{xy} = \frac{N}{dx dy dz} \frac{dz}{dt} = nu_z,$$

where u_z is the z component of fluid velocity. In general

$$\bar{\Gamma} = n\mathbf{u}.$$

The current density \mathbf{j} of the fluid parcel is the charge flux

$$\mathbf{j} = nq\mathbf{u}.$$

The last fluid variable we need to consider is the pressure, which in fact is the flux of momentum evaluated in the frame moving with \mathbf{u} .

1.2 Continuity equation

$$\frac{\partial}{\partial t} m = \frac{\partial}{\partial t} (\rho dx dy dz) = \sum \text{inward-flow} = \rho u_x dy dz$$

6 sides:

$$\int \frac{\partial \rho}{\partial t} dx^3 = - \oint \rho \mathbf{u} dA$$

divergence theorem: $\oint \mathbf{F} dA = \int (\nabla \mathbf{F}) dx^3$:

$$\int \frac{\partial \rho}{\partial t} dx^3 = - \int \nabla(\rho \mathbf{u}) dx^3$$

or

$$\int \left[\frac{\partial \rho}{\partial t} + \nabla(\rho \mathbf{u}) \right] dx^3 = 0$$

Continuity equation:

$$\boxed{\frac{\partial \rho}{\partial t} + \nabla(\rho \mathbf{u}) = 0}$$

In general, if \mathbf{H} is conserved, then

$$\frac{\partial}{\partial t} \mathbf{H} + \nabla(\mathbf{H}\mathbf{u}) = 0,$$

where $(\mathbf{H}\mathbf{u})$ is the flux of \mathbf{H} . If there is a source term $S = \frac{\text{new mass}}{dx^3}$, then we need to change the continuity equation to

$$\boxed{\frac{\partial \rho}{\partial t} + \nabla(\rho\mathbf{u}) = S.}$$